

**Erratum: Comparison of search templates for gravitational waves from binary inspiral**  
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Recently [1], an erratum has appeared correcting some coefficients in the computation of tails in the flux of gravitational waves  $\mathcal{L}$  from compact binaries in Ref. [2]. As a consequence, some post-Newtonian coefficients in the paper [3] are modified. The correction affects only the  $\eta$ -dependent terms in the coefficients at 2.5PN order, i.e. column 6 of Table I.

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TABLE . Taylor coefficients of the flux, phase, time and frequency.  $N$  denotes the “Newtonian value” and  $\theta = [\eta(t_{\text{iso}} - t)/(5m)]^{-1/8}$ . In all cases the  $k = 0$  coefficient is 1 and the  $k = 1$  coefficient is zero. In certain cases the 2.5 PN term involves  $v^5 \log v$  or  $\theta^5 \log \theta$  term rather than a  $v^5$  or  $\theta^5$  term. In those cases we conventionally include the  $\log v$  dependence in the listed coefficient. Chirp parameters  $\tau_k$ ,  $k \geq 1$ , are related to the expansion parameters  $t_k^v$  and  $\phi_k^v$  via  $\tau_k = (8\phi_k^v - 5t_k^v)/3$ . We have given the simplified expressions for these in all cases, except  $k = 5$  where no simplification occurs due to the presence of the log term in  $\phi_5^v$ .

$k$	$N$	2	3	4	5
$\hat{\mathcal{F}}_k$	$\frac{32\eta^2 v^{10}}{5}$	$-\frac{1247}{336} - \frac{35\eta}{12}$	$4\pi$	$-\frac{44711}{9072} + \frac{9271\eta}{504} + \frac{65\eta^2}{18}$	$-\left(\frac{8191}{672} + \frac{583\eta}{24}\right)\pi$
$\hat{t}_k^v$	$-\frac{5m}{256\eta v^8}$	$\frac{743}{252} + \frac{11\eta}{3}$	$-\frac{32\pi}{5}$	$\frac{3058673}{508032} + \frac{5429\eta}{504} + \frac{617\eta^2}{72}$	$-\left(\frac{7729}{252} - \frac{13}{3}\eta\right)\pi$
$\hat{\phi}_k^v$	$-\frac{1}{16\eta v^5}$	$\frac{3715}{1008} + \frac{55\eta}{12}$	$-10\pi$	$\frac{15293365}{1016064} + \frac{27145\eta}{1008} + \frac{3085\eta^2}{144}$	$\left(\frac{38645}{672} - \frac{65\eta}{8}\right)\pi \ln\left(\frac{v}{v_{\text{iso}}}\right)$
$\hat{\phi}_k^t$	$-\frac{2}{\eta\theta^5}$	$\frac{3715}{8064} + \frac{55\eta}{96}$	$-\frac{3\pi}{4}$	$\frac{9275495}{14450688} + \frac{284875\eta}{258048} + \frac{1855\eta^2}{2048}$	$\left(\frac{38645}{21504} - \frac{65\eta}{256}\right)\pi \ln\left(\frac{\theta}{\theta_{\text{iso}}}\right)$
$\hat{F}_k^t$	$\frac{\theta^3}{8\pi m}$	$\frac{743}{2688} + \frac{11\eta}{32}$	$-\frac{3\pi}{10}$	$\frac{1855099}{14450688} + \frac{56975\eta}{258048} + \frac{371\eta^2}{2048}$	$-\left(\frac{7729}{21504} - \frac{13}{256}\eta\right)\pi$
$\hat{\tau}_k$	$\frac{3}{128\eta}$	$\frac{5}{9}\left(\frac{743}{84} + 11\eta\right)$	$-16\pi$	$2\phi_4^v$	$\frac{1}{3}(8\phi_5^v - 5t_5^v)$

[1] L. Blanchet, Phys. Rev. D **71**, 129904(E) (2005).

[2] L. Blanchet, Phys. Rev. D **54**, 1417 (1996).

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